

# Checking for undesired vacua quickly at the 1-loop level

## Introducing Vevacious

Ben O'Leary  
in collaboration with  
José Eliel Camargo Molina, Werner Porod, and Florian Staub

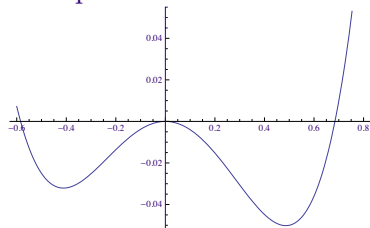
Julius-Maximilians-Universität Würzburg

Higgs parallel session,  
SUSY2013, ICTP Trieste,  
August 26th, 2013

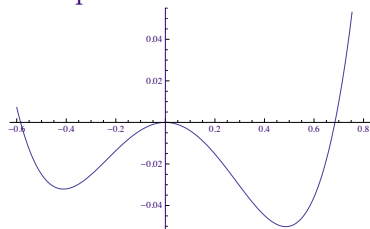


QFT potentials typically have multiple minima

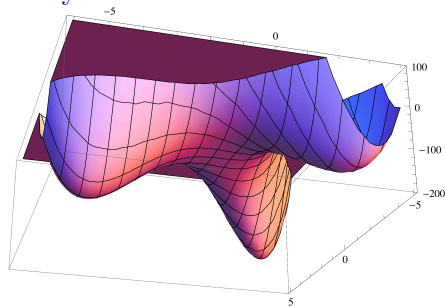
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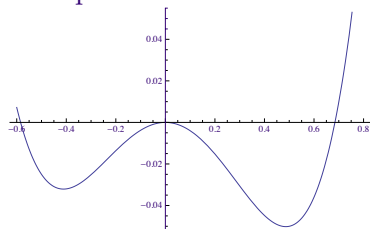
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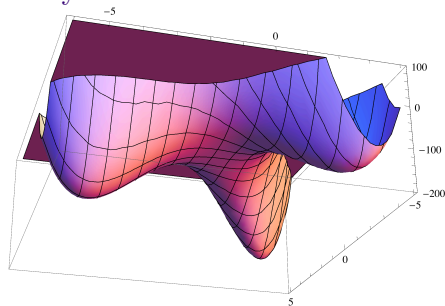


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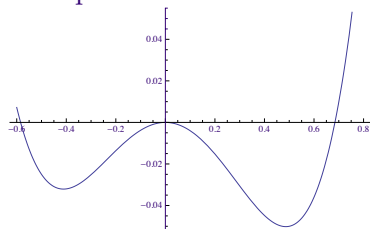


Finding global minimum not trivial!

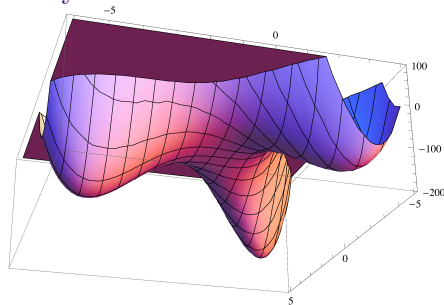
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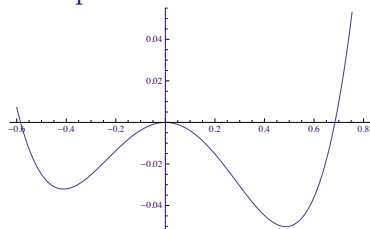
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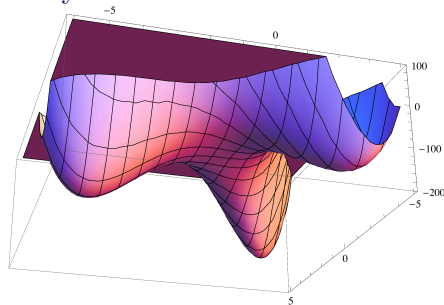
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Finding global minimum not trivial!

- ▶ Charge- and/or color-breaking (CCB) minima ( $\tilde{\tau}, \tilde{t}$  VEVs)?
- ▶ Desired VEV combination may not be global minimum (even non-CCB: NMSSM)





Gröbner bases:

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- ▶ Decomposition of system using fancy algebra
- ▶ Has been used to investigate NMSSM  
(Maniatis, von Manteuffel, Nachtmann, arXiv:hep-ph/0608314, EJPC)
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v e v a c i o u s

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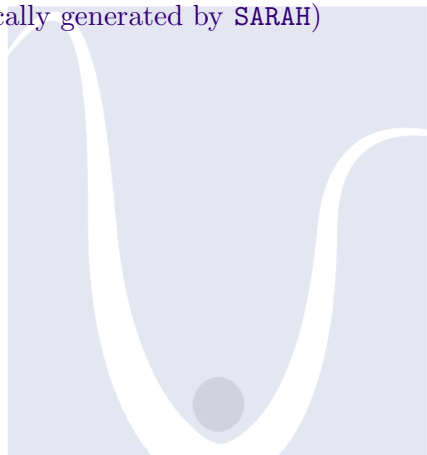
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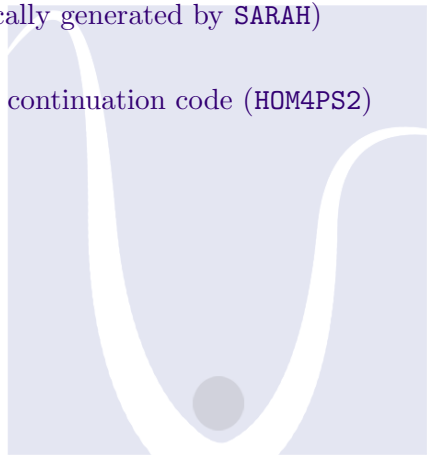
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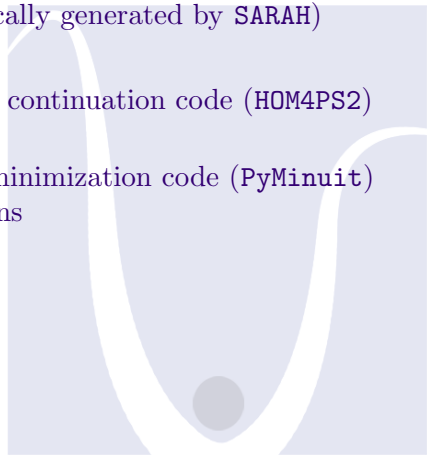
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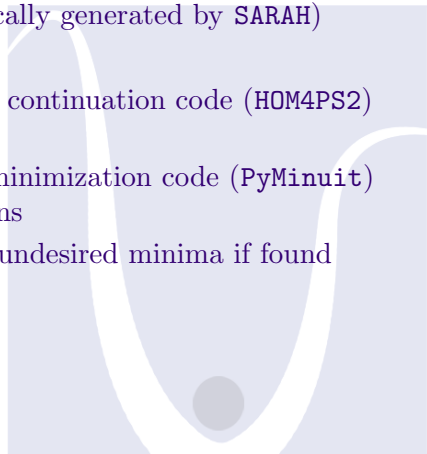
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Fast enough for scans! MSSM with additional non-zero VEVs for  $\tilde{\tau}_L, \tilde{\tau}_R, \tilde{t}_L, \tilde{t}_R$ : global minimum found within 5s on my laptop. (Tunneling time calculation varies: less than a second, up to 10 minutes.)

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<http://vevacious.hepforge.org/>



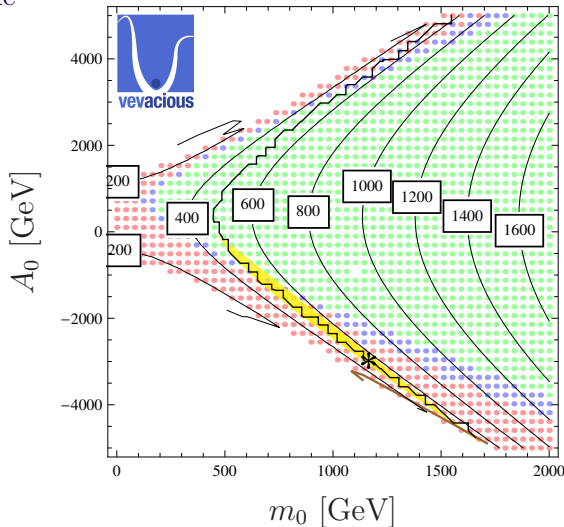


$M_{1/2} = 1110$  GeV,  $\tan \beta = 39.3$ ,  $\mu > 0$ ;  $m_{\tilde{\tau}_1}$  (GeV) contours  
 red: short-lived metastable ( $\tau_{\text{tunnel}} < 1.4$  Gy)  
 blue: long-lived metastable  
 green: stable

yellow region:  
 correct relic density

star:  
 best-fit point of  
[arXiv:1204.4199](https://arxiv.org/abs/1204.4199)  
 (Fittino)

wiggly line through  
 yellow:  
 neutralino LSP border



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- ▶ Estimating tunneling time strongly depends on relative depth and location of global minimum compared to input minimum:  $15s$  typical,  $500s$  for borderline cases

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V^{\text{tree}} = & \\
& \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u + \\
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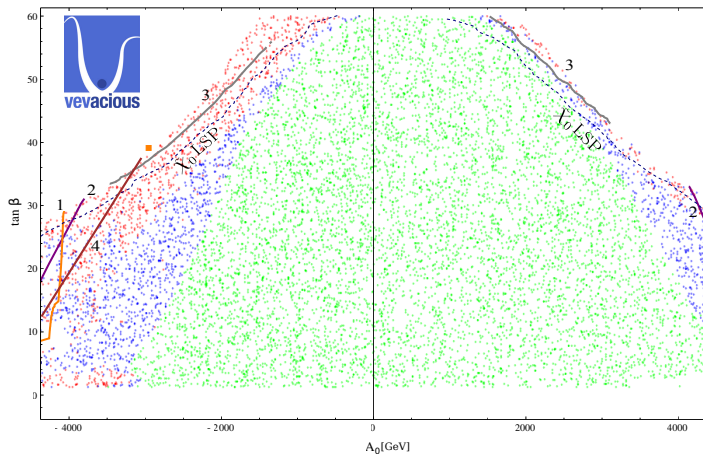
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(“ $A_\tau$ ”, “ $A_t$ ”: L. Alvarez-Gaumé, J. Polchinski, M. Wise, Nucl. Phys. B221;

“numeric”: Kitahara, Yoshinaga, arXiv:1303.0461, JHEP)

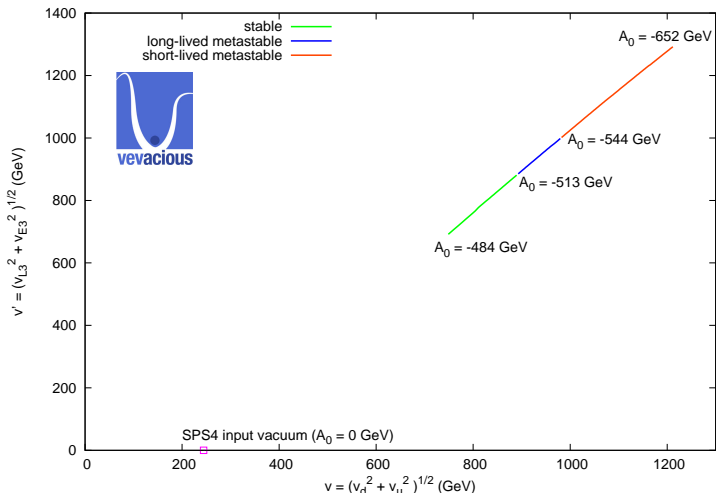


$$M_{1/2} = 1000 \text{ GeV}, m_0 = 1000 \text{ GeV}, \mu > 0$$



Purple: “ $A_\tau$ ”; Orange: “ $A_t$ ”; Grey: “numeric”  
 (Dark red: variant on “ $A_\tau$ ”)





$$m_0 = 400 \text{ GeV}, M_{1/2} = 300 \text{ GeV}, \tan \beta = 50, \mu > 0$$



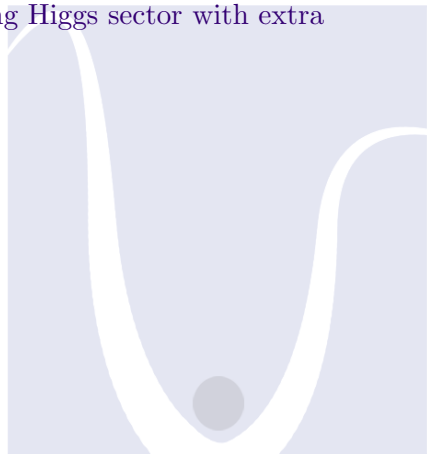
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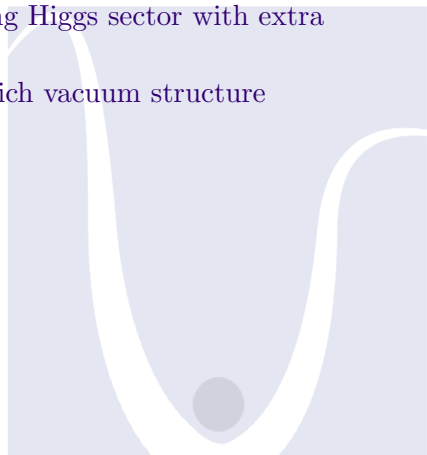
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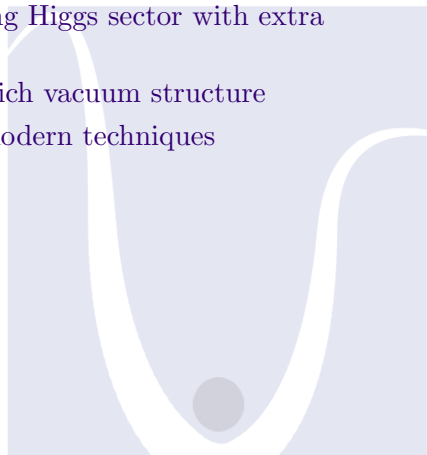
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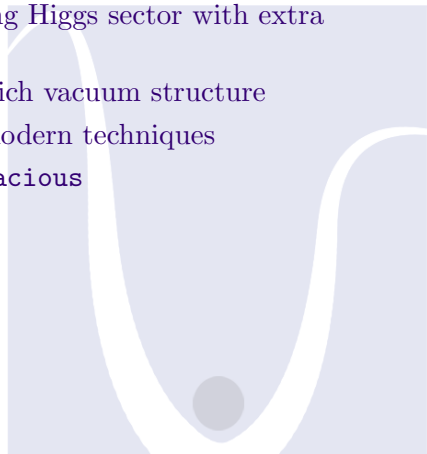
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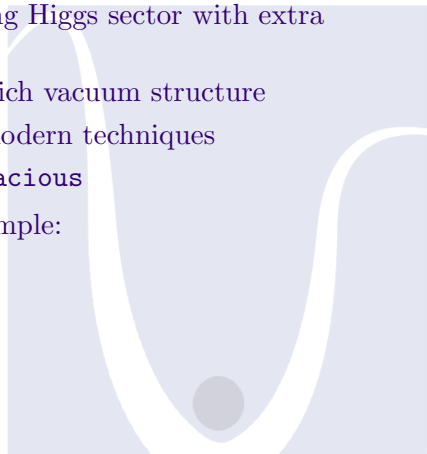


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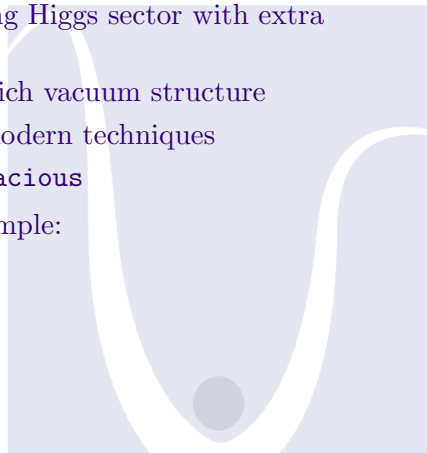
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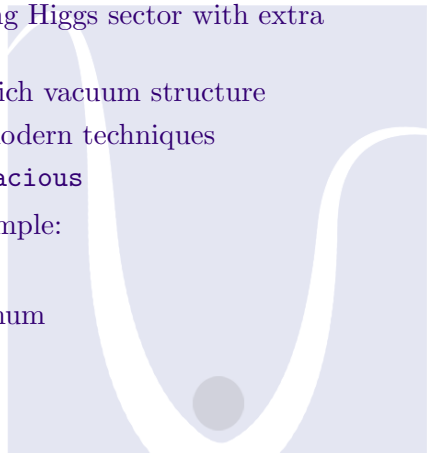
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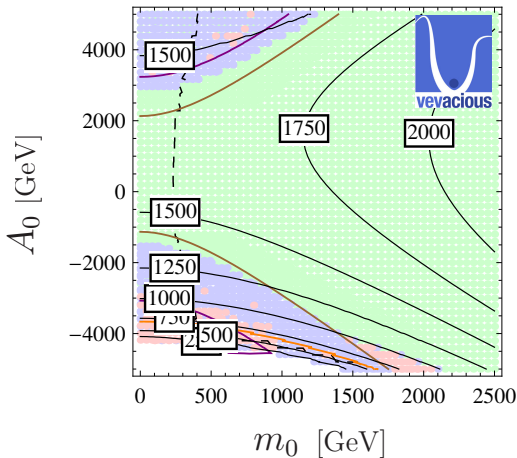
Backup slides

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orange line:  
 $A_t$  condition border

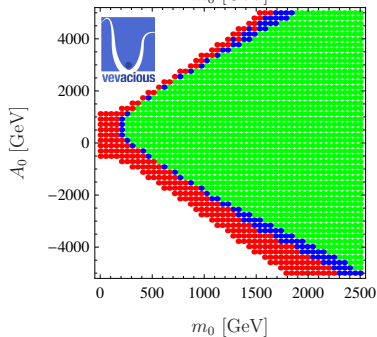
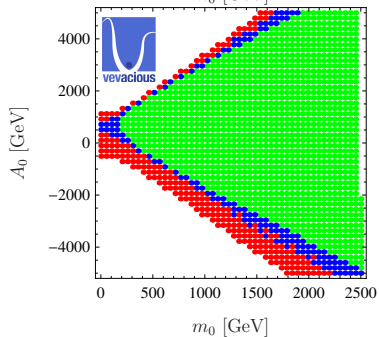
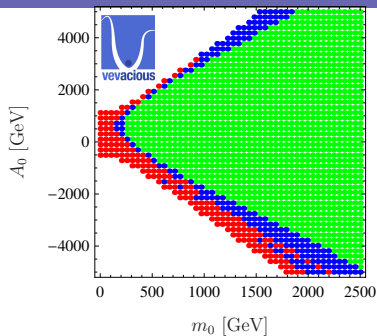
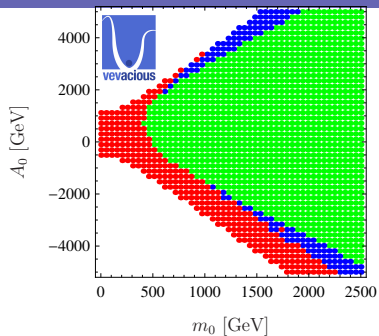
purple line:  
 $A_\tau$  condition border

dashed black line:  
neutralino LSP border



- ▶  $\Gamma / \text{volume} = Ae^{-B/\hbar}(1 + \mathcal{O}(\hbar))$
- ▶  $A$  is solitonic solution, should be  $\sim$  energy scale of potential
- ▶  $B \sim ([\text{surface tension}]/[\text{energy density difference}])^3$
- ▶ typically TeV-scale energy barriers, energy depth differences  $\Rightarrow$  roughly tunneling times of (factors of  $16\pi^2$  etc.)/TeV  $\ll$  age of Universe

# Scale and loop order dependence: halving $Q$



# Scale and loop order dependence: doubling $Q$

